

**Ch 22 - Sun and Its Solar System**

**Lesson I - The Sun**

Studying the sun should be done with caution. Observing the sun directly, especially with a telescope, can cause permanent eye damage. A \_\_\_\_\_ is an instrument that's used to determine the sun's chemical composition, temperature, and internal pressure. Solar physicists also use a \_\_\_\_\_, which projects a large image of the sun into a dark underground room. OSO-1 through OSO-8, Helios A & B, and Solar Max, are a few of the \_\_\_\_\_ used to study the sun and its radiation without interference from the Earth's atmosphere.

Though the sun is an average-sized star, it is much larger than earth. The diameter of the sun is \_\_\_\_\_ times Earth's diameter. The sun's volume can hold more than \_\_\_\_\_ Earth's. The sun is \_\_\_\_\_ kilometers from earth. It takes light from the sun approximately \_\_\_\_\_ minutes to reach earth

The three regions of the sun's atmosphere include the \_\_\_\_\_ (bright yellow; lower & denser part of the atmosphere), \_\_\_\_\_ (lower part of the outer atmosphere; colored red by glowing hydrogen), and the \_\_\_\_\_ (outermost portion of sun's atmosphere; faint, pearly light seen during total eclipse).

Dark spots on the photosphere, called \_\_\_\_\_, vary in size (larger than Earth's diameter), duration (from a few hours to a few months), and numbers (from none to over 100). These dark features appear in magnetic pairs (N-S) and may be 1500°C \_\_\_\_\_ than the surrounding photosphere. Studying their motion shows the sun rotates from \_\_\_\_\_ to \_\_\_\_\_. Because the sun is made of gases, its rate of rotation varies, about 25 days at the \_\_\_\_\_ and 36 days (NASA, 2007) at the \_\_\_\_\_.

The corona gives off a constant stream of electrically charged particles called \_\_\_\_\_. Solar events causing huge gusts of charged particles include \_\_\_\_\_ holes and \_\_\_\_\_ flares. As the charged particles from the sun interact with the Earth's magnetic field, it causes \_\_\_\_\_ (northern lights) and possible \_\_\_\_\_, which can disrupt radio signals and electrical service.

The sun's energy comes from the \_\_\_\_\_ of lighter elements into heavier ones. Einstein's famous equation,  $E=mc^2$  states that \_\_\_\_\_ can be converted into \_\_\_\_\_ and vice versa. The sun is mostly \_\_\_\_\_ and has enough matter to keep releasing energy for another \_\_\_\_\_ billion years.

**Lesson II - Observing the Solar System**

The solar system includes all objects that \_\_\_\_\_ (revolve) around the sun. Includes eight (8) planets [+dwarf planets], numerous \_\_\_\_\_ (moons), thousands of \_\_\_\_\_, millions of \_\_\_\_\_. Some orbits are nearly circular, others are highly elongated. Some objects travel close to the sun, others are billions of kilometers away. In the region of the solar system beyond Neptune are objects of the \_\_\_\_\_. Farther out in space is thought to exist the \_\_\_\_\_, possible source of long-period comets.

Due to the distance of stars from Earth, their positions relative to other stars do not appear to change. The position of planets, however, appear to move eastward in front of the stars; periodic backward (westward) loops called \_\_\_\_\_ motion occurs as earth passes another planet in its orbit

\_\_\_\_\_ (100 -165 AD) – Greek astronomer

Developed an earth-centered (\_\_\_\_\_) model of the solar system to predict the location of the planets; He imagined planets on small orbits (\_\_\_\_\_), and that the center of each epicycle moved around earth on a larger orbit (\_\_\_\_\_)

\_\_\_\_\_ – Polish astronomer

Proposed the \_\_\_\_\_ solar system – the Earth and other planets revolve around the sun

### Lesson III - Motion in the Solar System

\_\_\_\_\_ (1546-1601) – Danish

Built an astronomical observatory on an island; first long-term (20 years) observations of the sky; best observations made before telescope

\_\_\_\_\_ (1571-1630) – German

Used Tycho Brahe's data to develop **three laws of** \_\_\_\_\_

First Law – \_\_\_\_\_ **law** – every planet follows an elliptical orbit around the sun; the sun is located at one \_\_\_\_\_; the planet's distance from sun varies along its path (the point farthest from sun is \_\_\_\_\_, the point nearest is \_\_\_\_\_)

Second Law – \_\_\_\_\_ **law** – planets speed around sun varies; an imaginary line from the center of the sun to the center of a planet sweeps out the same area in a given time; planets move faster when they are closer to the sun

Third Law – \_\_\_\_\_ **law** – states that the period (P) of a planet squared is proportionate to the cube of its distance (D);  $P^2 = D^3$ ; \_\_\_\_\_ is the time it takes for a planet to travel one orbit around the sun; the farther a planet is from the sun, the longer its period (due to larger orbit & slower motion)

\_\_\_\_\_ (1564-1642) – Italian

Believed to be the first astronomer to use a \_\_\_\_\_ to study the sky; observed craters and mountains on the moon; discovered four moons in orbit around \_\_\_\_\_, which helped confirm the heliocentric view of the solar system (proved not all objects in the sky rotated around the Earth)

\_\_\_\_\_ (1642-1727) – English

Identified the force that kept planets in motion around the sun – \_\_\_\_\_; Newton's **universal law of** \_\_\_\_\_ showed that the force of gravity between any two objects is directly related to the \_\_\_\_\_ of the two objects, and inversely related to the square of the \_\_\_\_\_ between the centers of the two objects. More simply, gravitational force is 1) greater between objects of greater mass, and 2) decreases as distance between the centers of the two objects increases

Universal Law of Gravitation helps explain ...

... the reason that planets closest to sun move \_\_\_\_\_; why the speed of the the planets change (move \_\_\_\_\_ near the sun, \_\_\_\_\_ farther from the sun); used to determine masses of planets (when their period is known); cause of tides (Earth's moon); the long orbits of comets; the \_\_\_\_\_ (minimum speed needed to escape the gravitational pull) of a planet, moon, or asteroid.